## GDC

## Real-Time Reflections in Marla and Beyond

Martin Sobek

Lead Rendering Engineer



Martin Sobek has been passionate about making games since 1992. Martin studied computer science at Masaryk University in Czech Rebublic with a specialization in computer graphics. He joined Illusion Softworks in 2007 and worked on 'Mafia II'. He then moved to Hangar 13 in California in 2013 and led the rendering team toward a successful release of Mafia III. Now he is lead rendering engineer at Hangar 13 Brno, Czech Republic.

## Mafia III overview

Open world, $3^{\text {rd }}$ person, action adventure Story driven, yet not linear
Set in 1968 New Bordeaux Released October, 2016
PS4, Xbox One, Windows, Mac OS

Mafia III is running on custom engine, which is an evolution of engine used in Mafia II.

## Agenda

Motivation
Existing solutions
Ray casting on GPU
Reflection rendering
Reflections on rough surfaces
Timings, Results, Conclusion
Future work

## Motivation

With PBR, reflections are an essential part of material shading
Having proper reflections is a major step towards photorealism

Not happy with any of the existing solutions


Obvious case - reflection from wet road


Doesn't even look wet without reflections.

Most of the surfaces are quite rough, reflections still play major role.


## Existing solutions

Screen-space tracing
PROS
Doesn't require content authoring
Good performance
Low memory cost

## CONS



Only captures what's on screen
$\rightarrow$ Lots of missing information (especially for high roughness)
$\rightarrow$ Unstable with movement (camera or dynamic objects)

## Existing solutions

Pre-filtered cube-map look-up PROS

Simple to implement
Great performance
CONS
Floating reflections
Problems with transitions between CMs
Iteration issues (if pre-rendered)
Missing dynamic objects
Isotropic


To achieve anisotropy, we would need to pre-filter the CM with multiple kernel configurations that would make it much less practical.

## Existing solutions

Combination of SSR + Pre-filtered cube-maps
PROS
Simple to implement
Good performance
CONS
Partially: Floating reflections
Partially: Missing dynamic objects
Isotropic


Problems with transitions between CMs
Development iteration issues (if pre-rendered, need to re-render every time scene changes)
Stability issues (with camera movement)

Bad issues around main character in $3^{\text {rd }}$ person games.

## Existing solutions

SSR + Parallax-corrected cube-maps (pre-filtered)
PROS
Good performance
No floating reflections
Better transitions between CMs
CONS
Only works well for environments with certain shapes
More content authoring (scene approximation)
Partially: missing dynamic objects
Isotropic
Iteration issues (if pre-rendered)

Multiple variants exist. E.g.:
Kevin Bjorke: sphere approximation
Bartosz Czuba: box approximation
Seb Lagarde: convex approximation

## Existing solutions

Cone tracing
PROS
No floating reflections
Dynamic objects can be included
Robust
Doesn't require authoring
CONS
Requires run-time scene voxelization (difficult to implement)
Huge memory requirements
High GPU cost (scene update, tracing)
Isotropic

## Existing solutions summary

None of the existing solutions fulfilled all requirements:
Stability with camera movement
Good performance and memory cost
Working seamlessly in all environments (indoor, city, landscape)
Reasonable content authoring cost
Real-time update (scene changes)

## Problem breakdown

Problem \#1
General GPU-friendly ray casting
Find ray intersection with scene
Achieve mirror reflections (roughness=0)
Problem \#2
Proper BRDF on all materials
What rays to cast?
How to process the results


## Ray casting on GPU

Mesh/BVH
Branching
Non-coherent memory accesses
How to compute shading?
Voxels
Memory heavy
Non-trivial implementation
Depth texture
GPU-friendly
Trivial implementation
Not perfect coverage of the space

Update on mesh/BVH: New API (DX12 DXR) and HW has been announced that is supposed to address some of the issues.

## Covering space with 2D projections

## Cube-map covers space perfectly from a given POV

6 2D views

## Add depth

Works well if ray start position is close to CM origin

## Efficiency decreases with distance from origin

Tracing height-fields seems to be the right direction for nowadays GPUs.
We like the small implementation cost (we already have 2D rasterization implemented), low memory footprint and good performance.

Cube-map placed in camera covers reflection on vast majority of the pixels on the screen. Has been proven on a prototype.

But can't render a cube-map every frame! Sparse updates (like 1 side every frame) would result in reflections popping and latency.

## Multiple cube-maps

Pick best CM for ray start position
Switch to a different CM when ray enters "shadow region"
Use cube-map array


We've got 3 manually placed cube-maps on the right image.
Ray starts tracing the green CM, at some point gets to shadow region, red CM takes over. Ray reaches area without any coverage (implausible result) and blue CM takes over to finally find a hit.

Cube-map array: to be able to run single tracing pass.

The more complex the environment is, the less efficient the CM coverage is. Would be terrible for fractals but works well for typical environment that we live in.

## Cube-maps placement <br> Hand-placed CMs

Indoors: 1 CM for each room/hallway
City: Crossroad and every about 50 m on straight roads
Landscapes: Sparsely placed CMs (approx. every 100x100 m) Automatic backup CMs

Automatically placed CM in camera, if no hand-placed CM is around
Mainly used during development

Manually placed CM is always better than the automatic backup probes. It was used on open water areas for example.

Cube-map coverage issues
Manual placement: Need good tools
Dynamic objects: costly update $\rightarrow$ rely on SSR Not all pixels are covered
Inconsistent resolution (depends on distance from CM origin)
Thin objects (rails, poles, signs, ...) interrupt rays

Thin objects create aforementioned shadow regions that interrupt ray tracing.

## Our cube-map set-up

8 active geometry CMs
1 sky CM
Resolution of each 512 px, full MIP chain
Can't pre-render CMs offline
Dynamic time of day and weather
If you can pre-render, don't need separate sky CM

CM array slightly larger to be able to prepare new CM.

## Cube-map rendering

Pre-compute max view distance offline (for each side)


Only consider objects in the pre-computed CM range for rendering the CM.

## Cube-map rendering

## Single CHull scene query for all sides <br> Use geometry shader to output to affected sides Limited feature set

Use lower LODs
Only render static objects (and static lights)
No post-FX
No sky (sky is rendered into separate CM, geometry cube-maps contain sky-flag in alpha channel)

No specular, no reflections, diffuse only
(need some approximation for metallic)
No fog/volumetric effects
No transparent objects
No AA

We want to submit as few draw-calls as possible. Many static objects are large (terrain, buildings) and intersect more than 1 cube-map view frustum (end up in more sides). So we collect all objects (for all sides) and then only test, which sides are affected (fill to CB from CPU). Submit just 1 draw-call that outputs the object into multiple sides using geometry shader.

We have learned that Geometry Shaders aren't the most optimal way of attacking multi-viewport rendering, however is supported on all our platforms and is least intrusive from the shader combinations point of view.

We are rendering simpler LODs - these don't have many vertices, so in the end this is not an issue and we will stick to this solution.

No specular in CMs: not only it's an optimization but it also dramatically reduces noise in the result - specular has high intensity and frequency. Having specular baked in CMs isn't correct either since specular is view dependent - reflection in a mirror has different specular.

## Cube-map updating

Sky CM
Update every few frames (clouds, ToD)
Geometry CMs
Update dynamic lighting regularly (round robin)
Cache G-Buffer and static lighting
Render new when better CM has been found

Because of dynamic time of day and moving clouds, we need to update sky CM very often (several times per second). Sun is considered dynamic light.

## Active cube-maps selection

Might differ per project
We use 8 closest to the player, with 2 special cases
At least one outdoor CM
Penalty in vertical axis to separate floors
Possible improvements
Use bounding boxes (in/out, distance)
Use occlusion queries
Pre-compute best CM set for volumes

Indoors are typically more populated with CMs, so if player is standing in front of indoor location, all 8 closest might be inside. Outdoor would have no CM at all, so we always force at least one outdoor.

## Reflection rendering

## Algorithm overview

Down-sample G-Buffer, apply NDF
Trace screen, output distance
Trace cube-maps, output distance \& index
Resolve to color
Upscale

## G-Buffer down-sampling and jittering

Can't afford tracing at full resolution
Trace at half resolution
Bilinear down-sample not recommended
Incorrect depth on edges
Lost detail in normal \& roughness buffers

## G-Buffer down-sampling and jittering

Detect depth discontinuities
If edge is detected, discard "minor samples"
Pick random sample (exploit temporal filter)
Jitter normal (apply NDF)
Output (all at half-res)
RT0: Depth
RT1: Jittered normal and roughness
RT2: Original normal and roughness

Random sample: we actually alternate pixels in $2 \times 2$ block

## Screen-space tracing

Trace screen-space depth
Output: traveled distance, "finished" flag
Stencil mask for "finished" flag

Traveled distance:


Stencil mask (white means finished):


# Best cube-map selection - CPU 

Generate 8 cube-map index chains
For each starting CM, estimate best 3 consecutive CMs Based on distance only
Output: 8 4-item CM chains

| 0 | 1 | 5 | 2 |
| :--- | :--- | :--- | :--- |
| 1 | 0 | 3 | 4 |
| 2 | 1 | 7 | 3 |
| 3 | 2 | 4 | 1 |
| 4 | 5 | 6 | 7 |
| 5 | 4 | 6 | 7 |
| 6 | 4 | 5 | 7 |
| 7 | 2 | 4 | 5 |

This is something to be improved. We currently only find 3 closets CMs to each CM. It doesn't even take visibility into account.

## Best cube-map selection - GPU

Select best starting CM per pixel Use stencil (unfinished pixels)
Start at SSR end position
Assign score to each of 8 active CMs
Output CM index with best score


Score per pixel is assigned based on:

- Visibility (is that pixel visible from CM origin?)
- Distance from CM origin
- Ray direction vs. origin $\rightarrow$ point vector
- $\quad$ CM fade value (when adding/removing CM)


## Cube-map tracing

2 passes based on roughness (HQ/LQ)
Start with SSR end point, using best CM
If tracing fails, switch to next CM in chain and continue
If all CMs fail, use fallback
Output traveled distance and CM idx (where hit was found)

Roughness > 0.1: 16 steps, 100 m, scale 1.17 - 1.25, 3 refine iterations
Roughness <= 0.1: 24 steps, 300 m, scale $1.18-1.22,4$ refine iterations

## Tracing fallback solution in Mafia III

## Black reflection

Mostly OK
Really bad on very reflective surfaces (water, metals)
Simple lookup of best CM
Very different results when best CM was changing
Eliminate popping using temporal filter

## Current tracing fallback solution

"Cocoon" cube-map depth MIP
Use 1 MIP (e.g. MIP\#4 - 32x32) to store very smooth approximation of space
Large blur kernel with MAX filter ignoring sky
Pushing thin geometry away
Removing all edges
Caps windows
Tracing never fails
Preserves space but removes details
Similar idea to parallax corrected cube-maps, automatically generated

## Cocoon MIP example

Original depth:


Cocoon depth MIP:
Cocoon depth MiP:

## Cocoon MIP example

Full tracing with fallback:


Note how the stairway, columns and flower-pot is pushed to the background but windows are still at their correct location.

Compare to simple look-up, where the windows would be on wrong place.

## Generating cocoon MIP

## Top-down pass

Build a MIP chain using MAX filter ignoring sky
If all (4) pixels are sky, result is sky, otherwise sky pixels are discarded

$$
71
$$



## Generating cocoon MIP <br> Bottom-up pass <br> Lower MIPs (lower than cocoon)

Replace sky pixels with weighted MAX of neighborhood from lower MIP

## Cocoon MIP

Replace all pixels with weighted MAX of neighborhood samples from lower MIP

## Upper MIPs (higher than cocoon)

Replace sky pixels with cocoon MIP sample

Caps windows/sky - also works as an optimization for rays ending up at sky. Instead of burning all steps towards sky, ray hits the sky proxy sooner.

Weighted MAX:

```
float pivotSample = SAMPLE_4D_LOD( srcTex, srcSampler, float4( dir, srcArrayIdx ), srcMip ).r;
float depth= 0;
for each sample
{
    float smp = SAMPLE_4D_LOD( srcTex, srcSampler, float4( vec, srcArrayIdx ), srcMip ).r;
    float weight = pow( dot( vec, dir ), specPow );
    float currVal = pivotSample + ( smp - pivotSample ) * weight;
    depth = max( depth, currVal );
}
```


## GDC

## Cube-map tracing optimizations

Use lower depth MIPs for higher roughness
Pre-compute internal volume (AABB/sphere/convex hull)
Run as async compute shader (lose stencil)

## Color resolve passes - inputs

From cube-map renderer
Geometry color cube-map array
Sky cube-map
From previous reflection passes (half-res)
Linear depth
Jittered normals
Stencil mask for SSR
Traveled distance (combined SSR \& CM)
CM idx (for non-SSR finished pixels)
From shading pass
Diffuse shading buffer (you don't want specular here)

When tracing is finished (got traveled distance, stencil mask, possibly CM index per pixel), it can be resolved to color using the mentioned inputs.

## Color resolve passes

Half-res passes
Resolve SSR color
Resolve CM color
Full-res passes
Upscale half-res resolved buffer, generate low-roughness stencil mask
Resolve SSR on low-roughness pixels
Resolve CM on low-roughness pixels

## Color resolve shaders

## Compute ray end position:

rayDir $=$-reflect( viewVector, surfaceJitteredNormal )
endPos $=$ worldPos + rayDir $*$ traveledDistance
Fetch sky CM
SSR only
Project end position to screen space
Fetch diffuse shading buffer (including sky)
CM only
Fetch cmIdx
endPos $-=$ cmCenter[cmIdx]
Fetch color CM[cmIdx]
color $+=$ sky color * ( 1 - color.a )
Compute fog blend factor
Lerp( color, sky color, fog factor )

A little hack to add fog to reflections (fog is included neither in CM nor is SS diffuse shading buffer): because we have volumetric fog, which is non-trivial to compute for other rays than from camera, we simply fade towards sky color - which in fact is fog integrated over long distance.

## Upscale

Inputs:
Half-res color
Half-res unjittered normals
Half-res depth
Full-res normals
Full-res depth
Outputs:
Full-res color (high roughness pixels)
Stencil mask
Picks 1 sample from half-res color that best matches full-res normal \& depth

## Reflections on rough surfaces

## Possible approaches

Using pre-filtered MIPs
Visible edges
Isotropic


Pre-filtered MIPs:


Edges explanation:


Check out the edge artifact and missing elongation on the left image.

Diagram shows, how two neighbor pixels rays end up in a completely different location in the CM, the results are vastly different. CM is prefiltered from the point of view of its origin, not from the point of view of reflecting pixel.

## Possible approaches

Screen-space blur
Leaking \& losing detail
Stability issues
Isotropic
Reference:


On rough surfaces, the kernel is really large - would be very costly for real-time. That's why MIPs are used, so the blur can't be depth/normal/roughness aware. Note the big loss of normal map detail but also how it leaks across edges.

## Possible approaches

Importance sampling
Noise vs. performance
Need hundreds of samples to get noise-free result


We are shooting 1, 8, 32, 128 rays for every 4 pixels (still tracing at half resolution).

## Mafia III approach

Combination of screen-space blur and importance sampling 50 \% SS blur
50 \% importance sampling
Trade-off between leaking and noise
Large blur kernel (up to 25 \% of screen)
Need to use MIPs
Can't be depth-aware

Compute approximate reflection cone angle.
Halve the angle and jitter normal within this cone.
Output the ray traveled distance along with the reflection color.
Build color MIP chain.
For each pixel, estimate the MIP level to be used, based on traveled distance.

## Current approach

Mix of all $3+$ some tricks
50 \% importance sampling
50 \% using pre-filtered MIPs (both SSR and CM)
5-sample BRDF-weighted screen-space blur
Modified sample distribution
Temporal filter
Math is based on Blinn-Phong (not converted to GGX yet)


Note the leaking and loss of normal map detail.


## Importance sampling vs. pre-filtering - 100:0



Compare several mixtures of importance sampling vs. pre-filtering. 100 \% importance sampling is our reference.

Importance sampling vs. pre-filtering - 75:25


Importance sampling vs. pre-filtering - 50:50


Importance sampling vs. pre-filtering - 25:75


- Lost elongation
- Visible Edges
- Less correct - some surfaces look a lot different


## Combining importance sampling with pre-filtering

NDF produces vectors with angle [0, $\pi / 2$ ) from normal
Find angle, where probability drops below threshold (in our case 0.1)
Ignore all vector beyond this angle
Split angle among NDF and pre-filering
Modify NDF to produce vectors [0, angle/2)
Compute cone base radius and MIP level for angle/2

We lose a bit of the tail by ignoring all vectors, where
"cos(angle)^specPow < 0.1 " but on the other hand that helps reducing the noise quite a bit.

## Combining importance sampling with pre-filtering <br> Example: <br> roughness $=0.5 \rightarrow$ specPow $=30$ <br> angle $=\operatorname{acos}($ threshold $1 /$ specPow $)=0.387$

Changing importance sampling vs. pre-filtering ratio:
More importance sampling $\rightarrow$ more noise
More pre-filtering $\rightarrow$ less anisotropy


Blue graph is target NDF. Red line is threshold (0.1). We ignore regions, where blue is below red. Compute corresponding (cone) angle. Half of the cone is delivered using NDF (green), second half using pre-filtering (yellow).

## Pre-filtering cube-maps

## We do it at run-time $\rightarrow$ needs to be fast

## Build regular MIP chain

Choose texel scale (in our case 3.5x)
Pre-filter individual MIPs

Simple English: once we know our cone angle, we find cube-map MIP, where cone base radius is texelScale texels ( 3.5 texels).
Setting texel scale to 1 would cause pre-filtering of only 1 texel -> no pre-filtering at all.
Setting texel scale too high would increase the cost of pre-filtering (you need to add more taps) but also force sampling of higher MIP levels, which will cost additional performance in resolve pass.

When playing with this, cross-check with reference (1000+ taps from upper MIPs or base level).

Found more advanced run-time pre-filtering later - want to have a look at that:
http://research.nvidia.com/publication/real-time-global-illumination-using-precomputed-light-field-probes

## Pre-filtering cube-maps

MIP pre-filtering (in our case 29 taps):
numpixels $=2$ mipdax * texelScale
angle $=$ atan ( numpixels $/$ cmSize $/ 2$ )
specPow $=\log _{\cos (\text { angle })}$ ) threshold
Computing MIP level in resolve shader:
angle = AngleFromSpecPow( specPow ) // see previous slides
radius $=\tan ($ angle $)$ * traveledDist
cmRadius = radius / length( hitPosCM ) / texelScale
numpixels $=\max (1$, cmsize $/ 2 *$ cmRadius $)$
mipLevel $=\log 2$ ( numpixels )

## Modified NDF

Input: 2 random values $[0,1$, uniform distribution
Default Phong distribution:
$\theta=\operatorname{acos}($ rnd $1 /$ (specPow +1 ) $)$
$\varphi=2 * \pi *$ rnd2
Half-angle:
halfAngle $=0.5 *$ AngleFromSpecPow( specPow )
minRnd $=\cos (\text { halfAngle })^{\text {specPow }+1}$
$\theta=\operatorname{acos}\left(\left(\operatorname{minRnd}+(1-\text { minRnd })^{*} \text { rnd }\right)^{1 /(\text { specPow+1) })}\right.$

We don't care about the PHI angle for now but want to modify THETA, to get only angle/2 instead of angle. We inverse the function, find minimum random value and then scale the input random value to be in range [minRnd,1). Don't clamp the value, it needs to be linear operation to preserver the relative probabilities.

## Combined BRDF comparison

Reference $=\cos ($ angle $)$ specPow
Result $=\int_{-h}{ }^{h} \cos (\operatorname{clamp}(x+$ angle, $-\pi / 2, \pi / 2))$ specPow $* \cos (x)$ halfAngSpecPow $d x$

"Result" is what you get, if you modify NDF to half angle and sample MIP corresponding to half-angle.
h - half-angle
halfAngSpecPow - specular power corresponging to half-angle
angle $=\operatorname{acos(threshold~}{ }^{1 / \text { specPow }}$ )
halfAngSpecPow $=\log _{\cos (0.5 *}$ angle ) threshold
It's not $100 \%$ the same but it's pretty close

## Modified NDF

Concentrate as much variance as possible to neighborhood
The best pattern we found was a "+" pattern - assign each pixel a value of 0-4
Every pixel has all 5 "classes" around that it can sample in blur pass
Map class ID to ray direction

$$
\varphi=2 * \pi *(0.2 * \text { rnd2 }+ \text { GetSSJitterPlus( ssPos, frameCounter ) ) }
$$

Shuffle temporarily
SS pixels: Hemisphere slices:


Pixel class ID from screen-space position and frame ID: float GetSSJitterPlus (in const uint2 ssPos, in const uint txaaframeCounter )
const uint SAMPLES_-_OFFSET $=2$;
const uint SAMPLES_COUNT $=5$
const uint sampleIdx $=($ ssPos.x + SAMPLES_Y_OFFSET $*$ ssPos. $y+$ txaaframeCounter $) \%$ SAMPLES_COUNT;
return 1.0 / SAMPLES_Count * sampleIdx;
$2^{\text {nd }}$ modification of NDF is to concentrate color variance to a small neighborhood, to be able to blur that in SS blur pass and remove the noise. The assumption is that rays going in similar direction are more likely to result in similar color and vice versa. Focus direction variance to neighbor pixels. We found that shifting "+" pattern works pretty well for this purpose.

Blue noise might be a good alternative. Will try that later and compare the results.

## Neighbor sample reuse

Sample depth and normal of 4 neighbors
Same pattern as pixel classification
Use unjittered normals
Compute weighted average
Center tap: 1
Depth/roughness discontinuity: 0
Evaluate BRDF otherwise

If all the pixels have the same roughness and normal (flat, rough surface), you can look at it as multiple (temporal) samples. Just average them (assuming there is no discontinuity).

If roughness is very different, we haven't found a way, how to combine these samples.

With changing normals, the BRDF using unjittered normal seems to be a good metric.
For very small roughness, we would have to consider also view vector divergence between neighbor pixels. Instead of that (extra cycles), we simply fade this blur out.

## Temporal filter

We use up to 15:1 previous frame blend ratio Reflections view dependent

Compute view vector divergence (previous vs. current frame)
Compute divergence threshold based on roughness
Mirror reflections: zero divergence threshold but no issues with noise!
Rough reflections: high divergence threshold but not so much view dependency!
Invisible in last frame (or discarded due to divergence)
Evaluate extra 4 samples in centers of neighbor "+" elements Effectively up to 25 samples ( $5 \times 5$ )


We use variance clamping for mirror reflections (roughness $=0$ ) and we gradually increase the clamping window with growing roughness. Variance clamping is fully disabled when roughness $>0.1$.

Extra 4 samples: look at it as separable blur. But instead of 2-pass horizontal/vertical, we do " + " and tilted " $x$ " that is sampling the neighbor "+" centers.

## Step-by-step recap - tracing

Down-sample G-Buffer depth, normal (add jitter), roughness to half-res buffers
Stencil mask based on roughness (different tracing quality for high/low roughness)
2-pass (high/low roughness) SSR trace outputting traveled distance and FIN flag
Stencil mask for SSR finished pixels
Best CM select
2-pass (high/low roughness) CM trace outputting traveled distance and CM idx

## Step-by-step recap - post-tracing

Resolve to color (SSR + CM)
Neighbor sample reuse (screen-space blur)
Temporal filter for high roughness
Depth \& normal aware upscale to full res
Resolve low roughness at full res (using half res traveled distance)
Temporal filter for low roughness (with variance clamping)

## Timings (1080p @ PS4)



| Down-sample G-Buffer | 0.25 |
| :--- | ---: |
| SSR trace | 0.55 |
| Select best starting CM | 0.25 |
| CM trace | 0.9 |
| Half-res resolve | 0.35 |
| SSR |  |
| CM | 0.1 |
| Half-res blur | 0.25 |
| Half-res temporal | 0.17 |
| Upscale | 0.1 |
| Resolve | 0.41 |
| Temporal | 0.22 |
| Sum | 0.1 |

Captured before porting to async CS. Slightly above budget of 3.0 ms .

## Results

All the screenshots have been captured using Mafia III assets and the new tech.

Note that the new tech has NOT been shipped in Mafia III.






## Conclusion

Stable reflections when camera/dynamic objects move
Reasonable amount of manual work
Little pre-compute (max view distance, inner volume)
Real-time on nowadays gaming hardware
Scalable in terms of:
Lighting changes: re-light cube-maps
Geometry changes (destruction): re-render affected cube-maps
Scene complexity: adjust amount of cube-maps

## Future work

Convert to GGX
Temporal re-projection using reflection depth
Improve upscaling pass
Pre-compute optimal starting CM and chain
Investigate automatic probe placement
Investigate better handling of off-screen dynamic objects

## Thanks

Petr Smílek

First implementation
Naty Hoffman
Consulting
Rinaldo Tjan
Testing and feedbacking
Tianli Bi
Optimizations
Jiří Štempin
Code support

## Eva Tajovská

Help with presentation
Jan Marvánek
Help with presentation
Radim Doleček
Help with presentation
Petr Záveský
Help with presentation
Sebastien Lagarde
Proof review

## References

Umenhoffer, Patow, Szirmay-Kalos. 2007. GPU Gems 3 - Chapter 17. Robust Multiple Specular Reflections and Refractions https://developer.nvidia.com/gpugems/GPUGems3/gpugems3 ch17.html
Stachowiak. SIGGRAPH2015. Stochastic Screen-Space Reflections
http:://advances.realtimerendering.com/s2015/Stochastic\%20Screen-Space\%20Reflections.pptx
Robinson, Shirley. 2009. Image-Space Gathering
http://www.nvidia.com/object/nvidia research pub 015.html
Valient. GDC2014. Taking Killzone Shadow Fall Image Quality into the Next Generation
https://www.guerrilla-games.com/read/taking-killzone-shadow-fall-image-quality-into-the-next-generation-1
Lagarde. SIGGRAPH2012. Parallax Corrected Cube-Maps
https://seblagarde.files.wordpress.com/2012/08/parallax corrected cubemap-siggraph2012.pdf
Bjorke. 2004. GPU Gems - Chapter 19. Image-Based Lighting
http://developer.download. nvidia.com/books/HTML/gpugems/gpugems ch19.html
Czuba. 2011. Box Projected Cube Environment Mapping
https://blenderartists.ora/forum/archive/index.php/t-209688.html
Manson, Sloan. 2016. Fast Filtering of Reflection Probes
https://www.ppsloan.org/publications/ggx filtering.pdf
McGuire, Mara, Nowrouzezahrai, Luebke. 2017. Real-Time Global Illumination using Precomputed Light Field Probes http://research.nvidia.com/publication/real-time-global-illumination-using-precomputed-light-field-probes

## .....questions?



## GOC

 GAME DEVELOPERS CONFERENCE $/$ MARCH 19-23, 2018 | EXPO: MARCH 21-23, 2018 \#GDC18
## Bonus slides

## Cube-map tracing pseudo-code

stepScale $=$ Rand ( stepScaleMin, stepScaleMax )
currStep $=$ ComputelnitialStep ( maxRayLength, stepScale
bestCMIdx = FetchBestCMIdx
bestCMIdx $=$ FetchBestC
currCMIdx $=$ bestCMIdx
currCMIdx $=$ be
usedCMs $=0$
currPos $-=$ cmCenter[currCMIdx] // currPos is always in CM space
for each step
currPos += currStep
cmDepth $=$ FetchCMDepth ( bilinear, currPos )
cmDepthPoint $=$ FetchCMDepth ( point, currPos )
cmDepth $=$ clamp ( cmDepth, cmDepthPoint - threshold, cmDepthPont + threshold )
cmDist = length (currPos ) // Note: sqrt can be avoided
if cmDist > AddBias( cmDepth )
if ComputeMassDepth ( currPos, currStep ) + cmDepth $>$ cmDist
// Hit has been found
numRefineSteps++
if numRefineSteps >= maxRefineSteps success = true succes
break
else
currPos $-=$ currStep currStep * $=0.5$


CM depth texture contains distance from CM origin instead of linear depth

- Simpler math
- Eliminate pre-filtering issues on CM edges

